

Design of the Calibration Unit for the MOAO Demonstrator RAVEN

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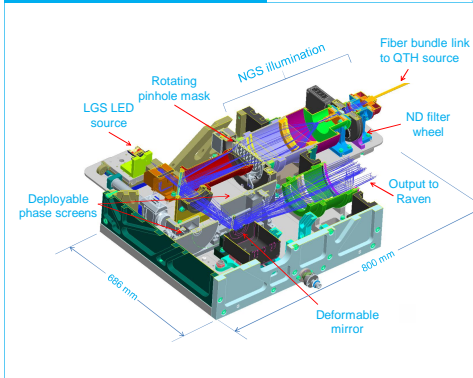
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ABSTRACT

The UVic AO Lab in collaboration with HIA and the Subaru telescope is currently designing Raven, a multi-object adaptive optics (MOAO) demonstrator that will be coupled to the Subaru Infrared Camera and Spectrograph (IRCS). Its main goal will be to demonstrate MOAO feasibility on the sky while allowing astronomers to benefit from the increased observing efficiency associated with such systems.

INO is responsible for the Raven calibration unit design and fabrication. This sub-system consists in a telescope simulator that will allow aligning Raven's components during its integration, testing its AO performances in the laboratory and at the telescope, and calibrating the AO system by building the interaction matrix and measuring the non-common path aberrations (NCPA). This sub-system is described in this poster.

SYSTEM OVERVIEW



SOURCES REQUIREMENTS

NGS parameters	
Source grid	Square grid with 9 x 9 sources (step of ~10.7 mm)
Rotation speed	Up to 10 ⁷ /s to simulate sky rotation
Waveband	0.6 – 1.8 μm Goal: 0.6 to 2.5 μm
Source width	Possibility to select diffraction limited (in the IR) or seeing limited (0.6 arcsec) sources
Star magnitude	< 10 to 16
LGS parameters	
Wavelength	589 nm
Source width	1 arcsec
Altitude conjugation	85 to 180 km
Star magnitude	6 to 11

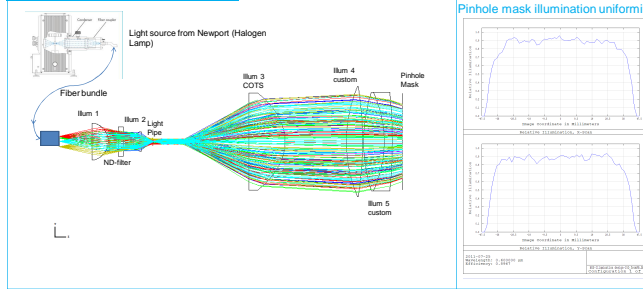
EXPECTED OPTICAL PERFORMANCES

Parameter	Requirement	Nominal value	With tolerances
NGS focal plane diameter in the Raven focal plane	84 mm Goal of 112 mm	Up to 86 mm	No significant change with tolerances
Focal plane curvature	< 2.3 m	< 2.3 m up to λ = 1.8 μm	No significant change with tolerances
Coma	< 1.3 μm	< 0.01 μm	No significant change with tolerances
Working F/#	13.6 +/- 0.3%	13.58 @ 1.1 μm	No significant change with tolerances
Rays position variation at exit pupil edge over full waveband at all FoV	1% Goal < 0.6%	1.1%	90% confidence: 1.2% 50% confidence: 1.1%
Edge FoV relative illumination vs the center FoV	> 50%	80% at radii of 35 mm	No significant change with tolerances
Image quality < 1 μm	RMS spot size < 100 μm	< 45 μm	No significant change with tolerances
Image quality > 1 μm	Diffraction limited (Strehl > 0.8)	Strehl > 0.90 for FoV < 64 mm	90% confidence: Strehl > 0.8 for FoV < 64 mm
Lateral color between 0.6 < A < 0.9 μm	≤ 100 μm for FoV < 64 μm	-2 μm for FoV < 64 mm -40 μm for FoV < 78 mm	No significant change with tolerances
Lateral color between 0.9 < A < 1.8 μm	≤ 20 μm for FoV < 64 μm	Max of ~ 15 μm	No significant change with tolerances
Axial color	< 580 μm	λ < 1.8 μm: 186 μm λ < 2.5 μm: 581 μm	90% confidence: 220 μm 50% confidence: 160 μm 90% confidence: 637 μm 50% confidence: 550 μm

INTRODUCTION

- Raven:
 - 3 natural guide stars (NGS), 1 laser guide star (LGS) and 2 science fields,
 - NGS wavefront sensing in R-band,
 - IRCS used science waveband: 0.9 to 2.5 μm.
- Raven CU:
 - Provides 9x9 NGS over a 2.7° FOV,
 - Provides an on-axis LGS,
 - Altitude conjugation: from 85 to 180 km,
 - Provides a bright on-axis alignment source,
 - 2 deployable phase screens:
 - Conjugated to 5 km and 10 km,
 - DM conjugated to ground level,
 - Remotely selectable NGS pinhole sizes:
 - Diffraction limited and seeing limited (0.6"),
 - Rotating NGS pinhole mask to simulate sky rotation,
 - Adjustable sources brightness,
 - Status: In fabrication,
 - Expected delivery: January 2012.

NGS ILLUMINATION



OPTICAL RELAY

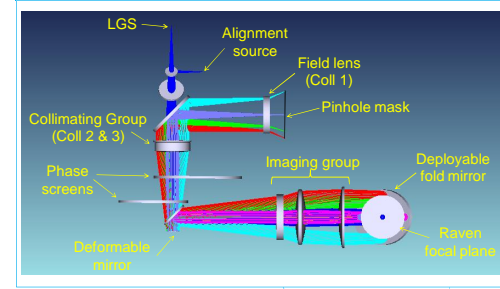
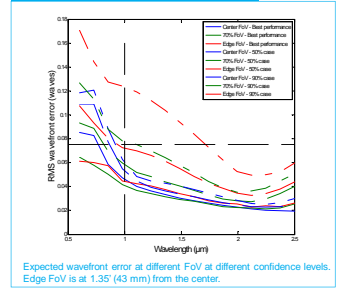


IMAGE QUALITY



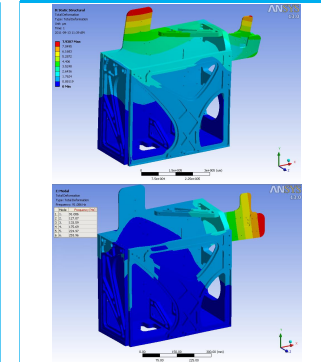
INO MECHANICAL TOLERANCE MODEL

- Goal:
 - Eliminate unnecessary adjustments to reduce risks and costs while achieving required positioning precision
- Methodology:
 - Replace linear and RSS sums of mechanical tolerances with Monte Carlo statistical model based on real statistical data
 - Update Zemax Monte Carlo model with realistic mechanical statistical values
 - Iteratively adjust tolerances to achieve optimal optical performances vs cost

TOLERANCES

Illumination Positioning Tolerances Requirements					Illumination Expected Positions Variations				
Group	Part	Thickness (mm)	Decenter (mm)	Tilt (°)	Group	Part	Thickness (mm)	Decenter (mm)	Tilt (°)
Group 1	FB	0.2	ADI	ADI	Group 1	FB	0.2	0.1	0.1
	Illum 1	0.5	0.2	0.2		Illum 1	0.5	0.1	0.1
	Filter NG	0.5	-	0.3		Filter NG	0.5	-	0.2
	Illum 2	0.2	0.2	0.2		Illum 2	ADI	0.1	0.1
	Light pipe	ADI	-	0.1		Light pipe	ADI	ADI	0.15
Group 2	Illum 3	0.1	0.2	0.1	Group 2	Illum 3	0.15	0.1	0.1
	Illum 4	0.25	0.2	0.2		Illum 4	0.1	0.1	0.1
	Illum 5	0.5	0.2	0.1		Illum 5	0.25	0.1	0.1
	PS 10 km	-	-	1		PS 10 km	1.000	-	1.000
	PS 5 km	-	-	1		PS 5 km	0.500	-	1.000
Group 3	BS	ADI	0.05	0.01	Group 3	BS	ADI	0.013	0.010
	DM	ADI	0.015	0.01		DM	ADI	ADI	ADI
	Imp1	0.25	0.015	0.01		Imp1	0.015	0.010	0.010
	Imp2	0.25	0.015	0.01		Imp2	0.015	0.010	0.010
	Imp3	0.25	0.015	0.01		Imp3	ADI	0.027	0.009
Group 4	M1	comp	-	-	Group 4	M1	comp	-	-
	Group 1A	Coll 1, 2 & 3	0.05	0.02		Group 1A	Coll 1, 2 & 3	0.01	0.018
	Group 1B	Coll 2 & 3	0.025	0.025		Group 1B	Coll 2 & 3	0.020	0.025
Group 2	Imp 1, 2 & 3	0.05	0.02	Group 2	Imp 1, 2 & 3	0.020	0.020		

FINITE ELEMENT ANALYSIS



METHODOLOGY

- Optical tolerances to meet requirements
 - Generation of 100 statistical Monte Carlo realizations including positioning and manufacturing errors on optical elements using Zemax
 - Expected parameters value with a 50% and 90% confidence are evaluated from the MC files
- Mechanical tolerances to meet optical tolerances
 - Lenses are aligned within tolerances using INO's Trioptic centering machine
 - Groups are aligned within tolerances by manufacturing tolerances (see INO mechanical tolerance model)
 - Finite element analysis to include deflections
- Adjustments: available on DM, on fiber bundle and on light pipe

CHALLENGES

- System with many optical interfaces with focal and pupil plane requirements.
- Control NGS chromaticity over a broadband (0.6 to 2.5 μm) in the focal and the exit pupil plane.
- Fast track project: 7 months from PO to delivery time.

SUMMARY

- All requirements are expected to be met when considering manufacturing and alignment tolerances.
- Small number of adjustments necessary.
- Expected delivery: January 2012