

# Numerical simulations of an Extreme AO system for an ELT

Miska Le Louarn, R. Clare, C. Bechet  
ESO, Garching

With help from Austrian AO Team & Frim3D Team

# Introduction

---

- Full end to end modeling (“Octopus”) of AO
  - Shifting phase screens
  - Diffractive WFS model
  - Measurements for each subap
  - Reconstructor
  - DM shape → residual phase
  - Closed loop
- Goals of the study:
  - Demonstrate that our full XAO model works and is tractable for a 42m telescope
  - Investigate XAO PSF as provided by the simulator
  - Compare reconstructors (“Austrian in-kind contribution to accession to ESO”)
    - This lead also to some MCAO results which will be presented as well

# XAO system parameters

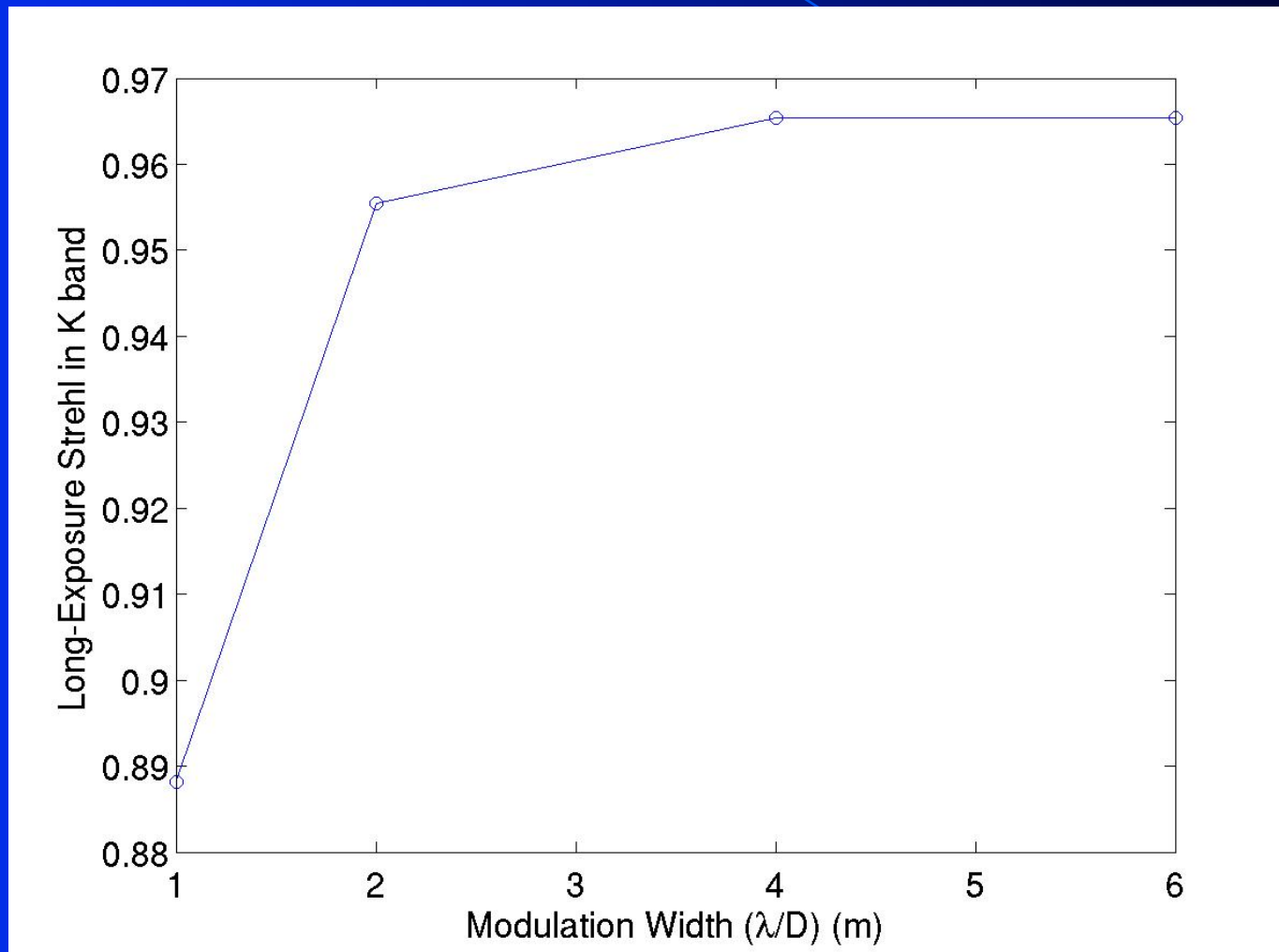
---

- 42m telescope, with central obstruction, no spiders, no segments, no wind shake.
- Pyramid sensor @700nm, with modulation
- 200x200 “sub-apertures” → ~20cm
- 3 kHz, 2 frames delay.
- Seeing: 0.8'' , Tau0: ~3ms
- No Woofer - tweeter: PYR sees all turbulence when loop is open (worst case for PYR linearity)
- Pupil: 2000 pixels ↔ 42m
- PSFs calculated at K-Band (unless otherwise noted)
- Static aberrations not considered, only “basic” atmospheric AO errors
- Temporal control: Simple integrator
- PYR module written by Ch. Verinaud

# Modulation

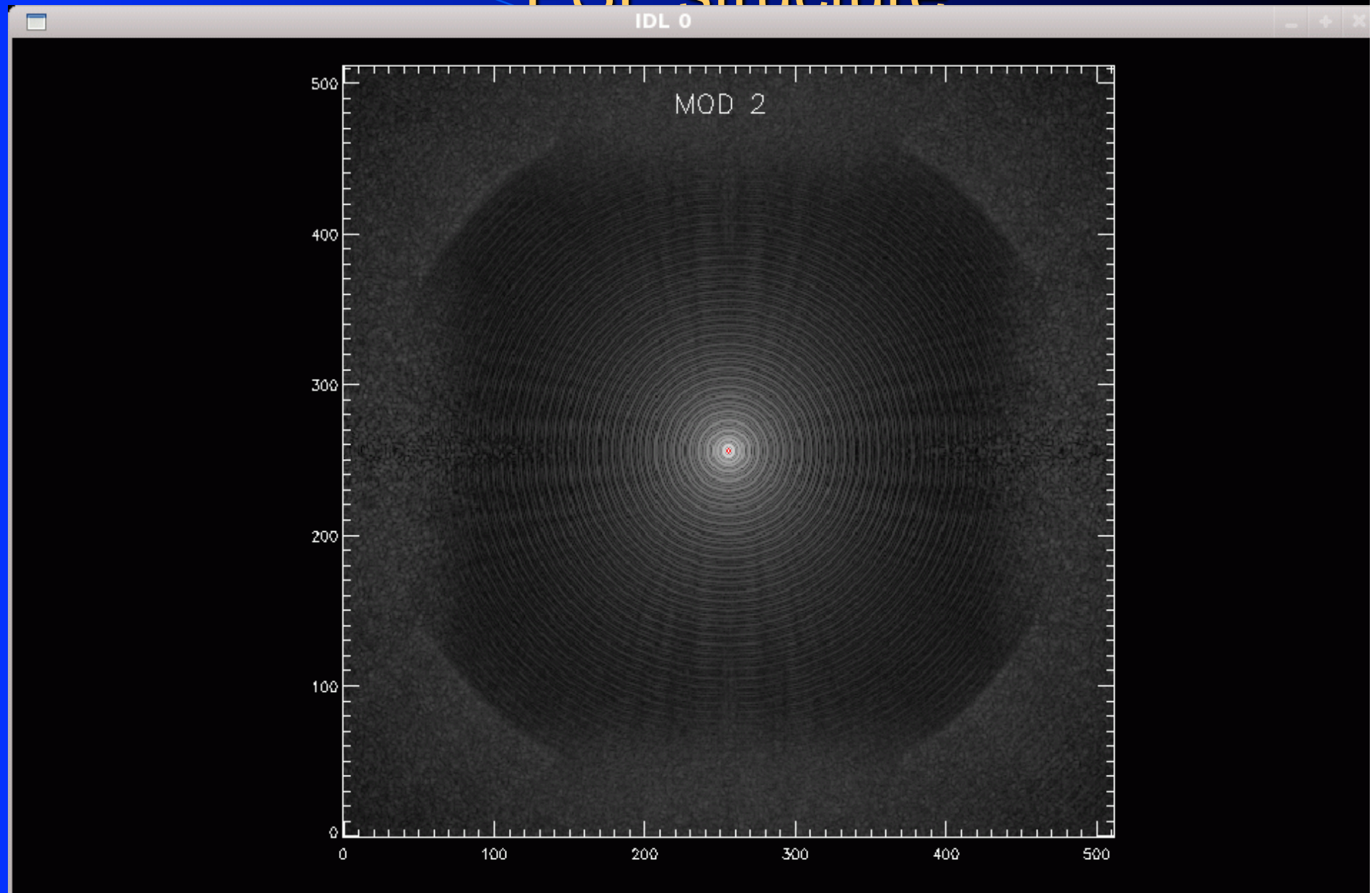
- Simulation tool allows to modulate PYR (square pattern)
- Larger modulation is more computation time intensive:
  - $2 \lambda / D$  of modulation  $\rightarrow$  16 points
  - $6 \lambda / D$  of modulation  $\rightarrow$  48 points
- Modulation is fully parallelized
- Still time consuming: 3h (mod 2) -7h (mod 6) for 500 iterations
- Allows to increase linearity range of PYR measurements
- Different modulations change behavior of PSF, even if Strehls are very similar

# Modulation & Strehl (high flux)



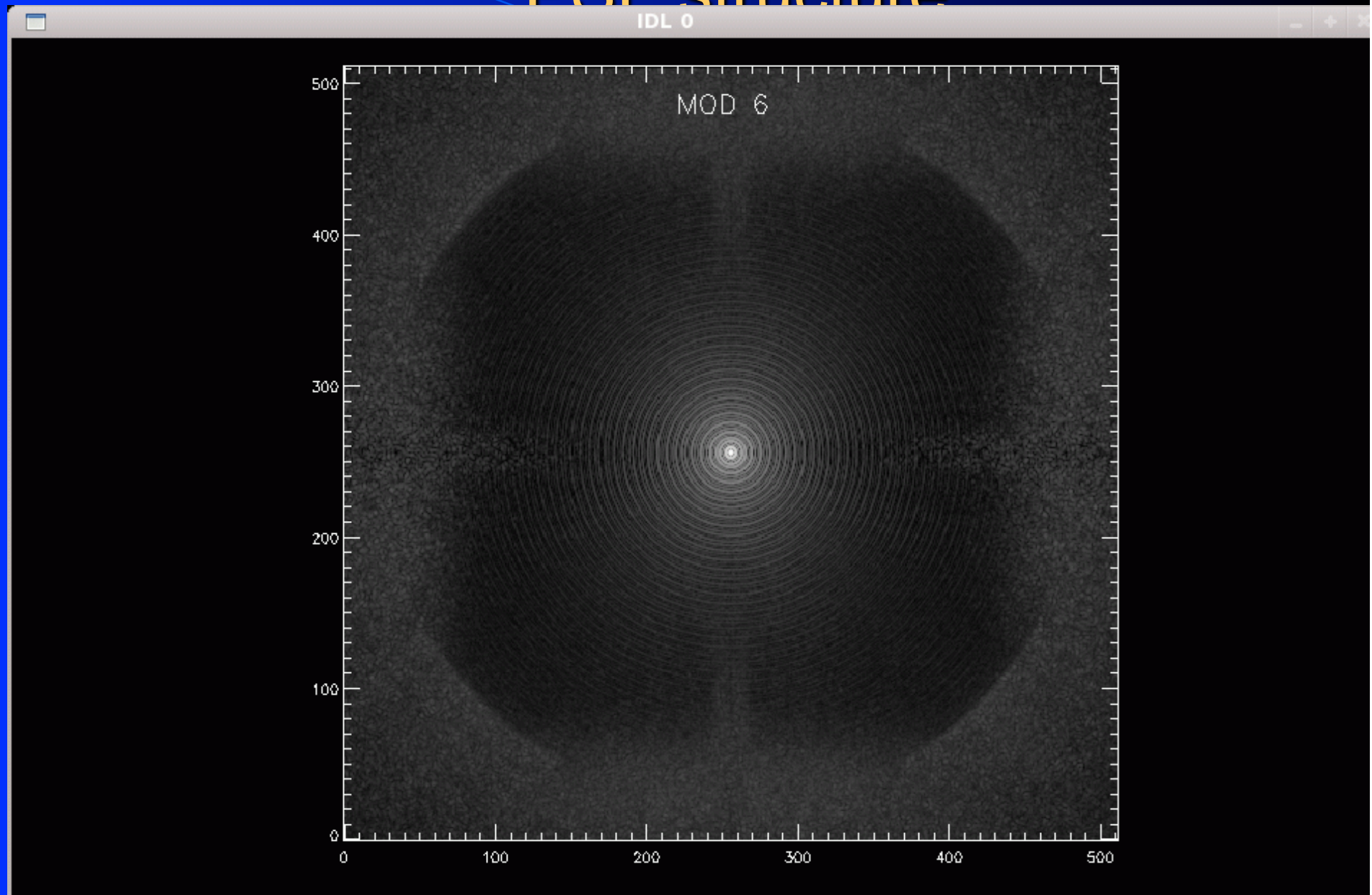
→ Modulation of 4 is chosen

# PSF structure



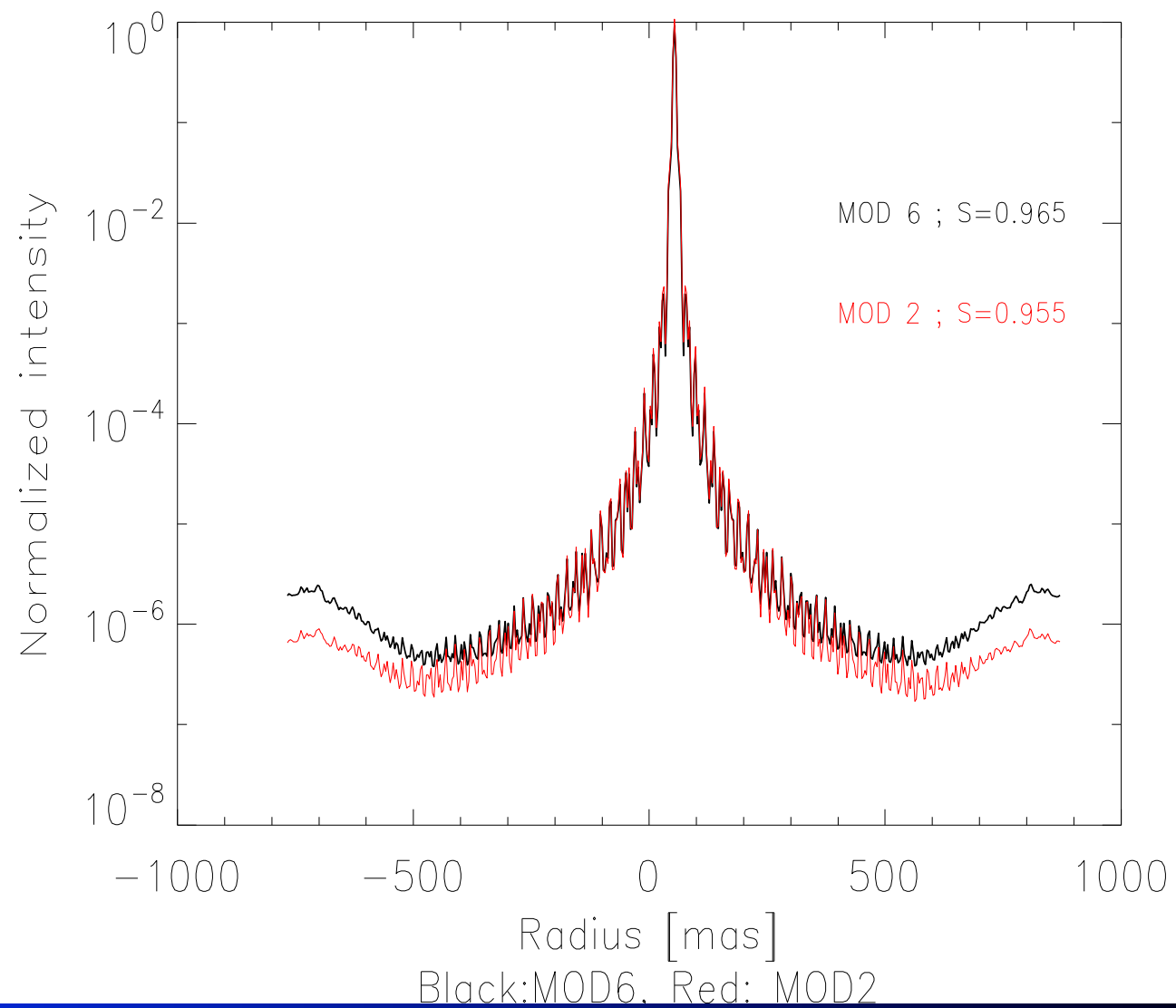


# PSF structure



Scale is simulation pixels (5.3 mas/pix)

# Impact of modulation at high flux



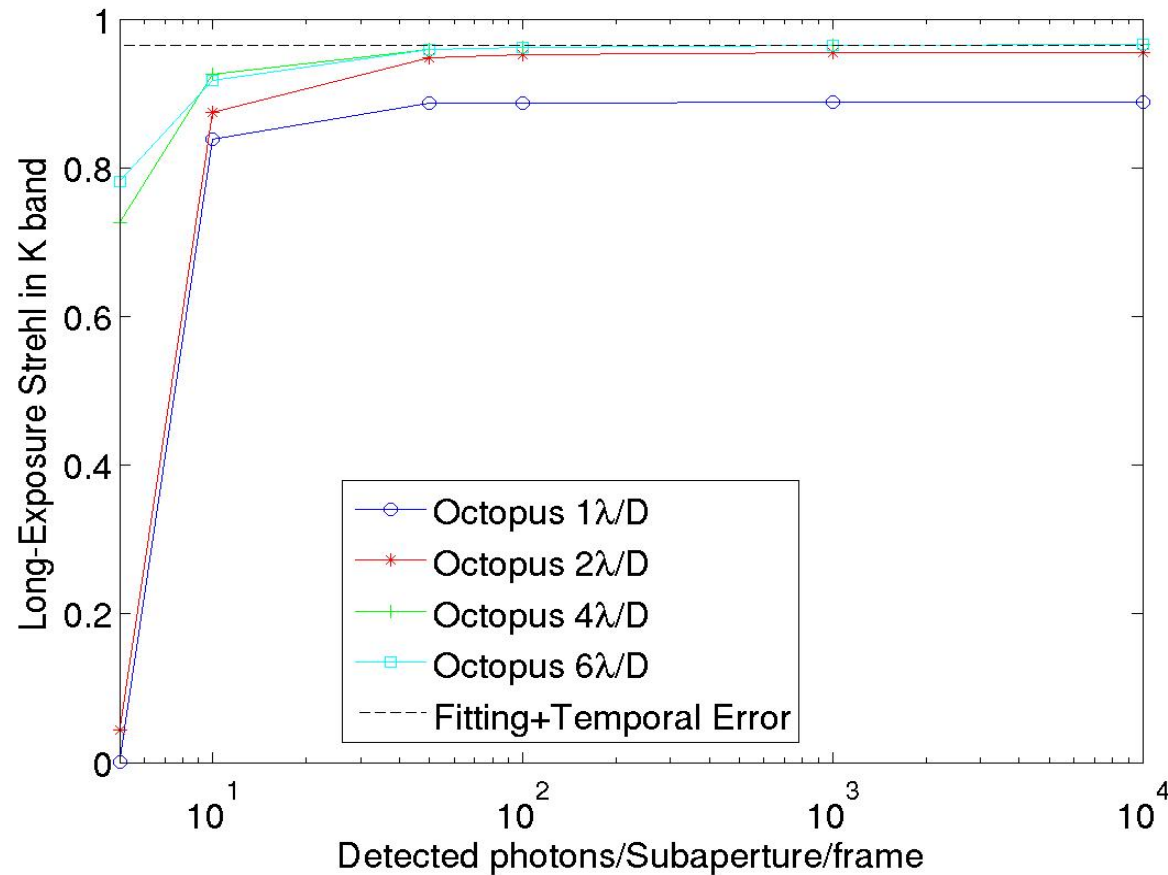


# Modulation & limiting magnitude

---

- Next we study the impact of modulation is on the limiting magnitude
- 2.8e of RON
- Optimize loop gain for each flux
- At low flux, amount of regularization when building command matrix is increased
- Framerate optimization not yet done (i.e. running slower to get less effect from RON).

# Photon flux



$d$  subap size

$V$  characteristic wind speed

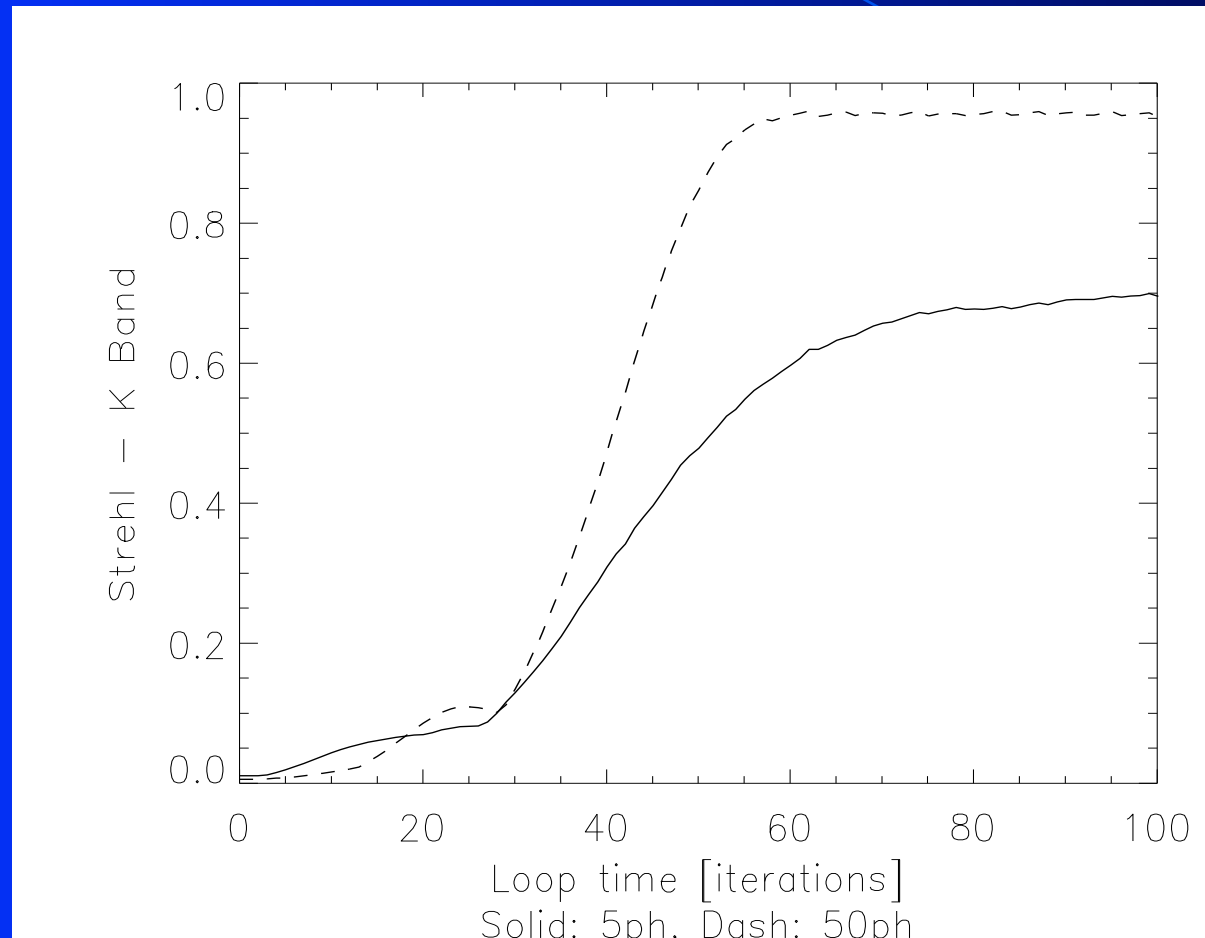
$T$  delay

$r_0$  the Fried parameter

$$\sigma_{fitting}^2 = 0.26(d/r_0)^{(5/3)},$$

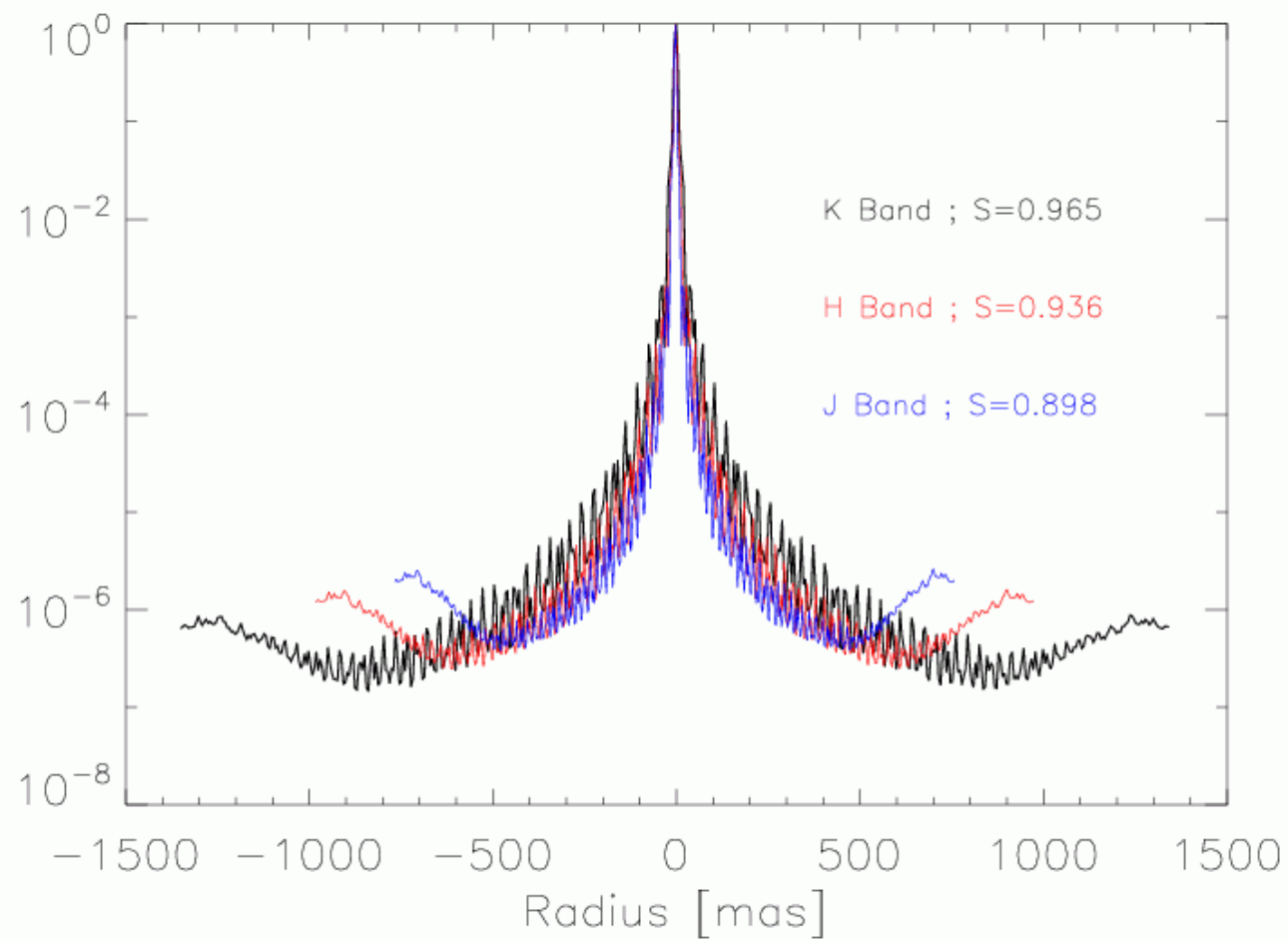
$$\sigma_{delay}^2 = 6.88(VT/r_0)^{(5/3)},$$

# Loop closing at low flux

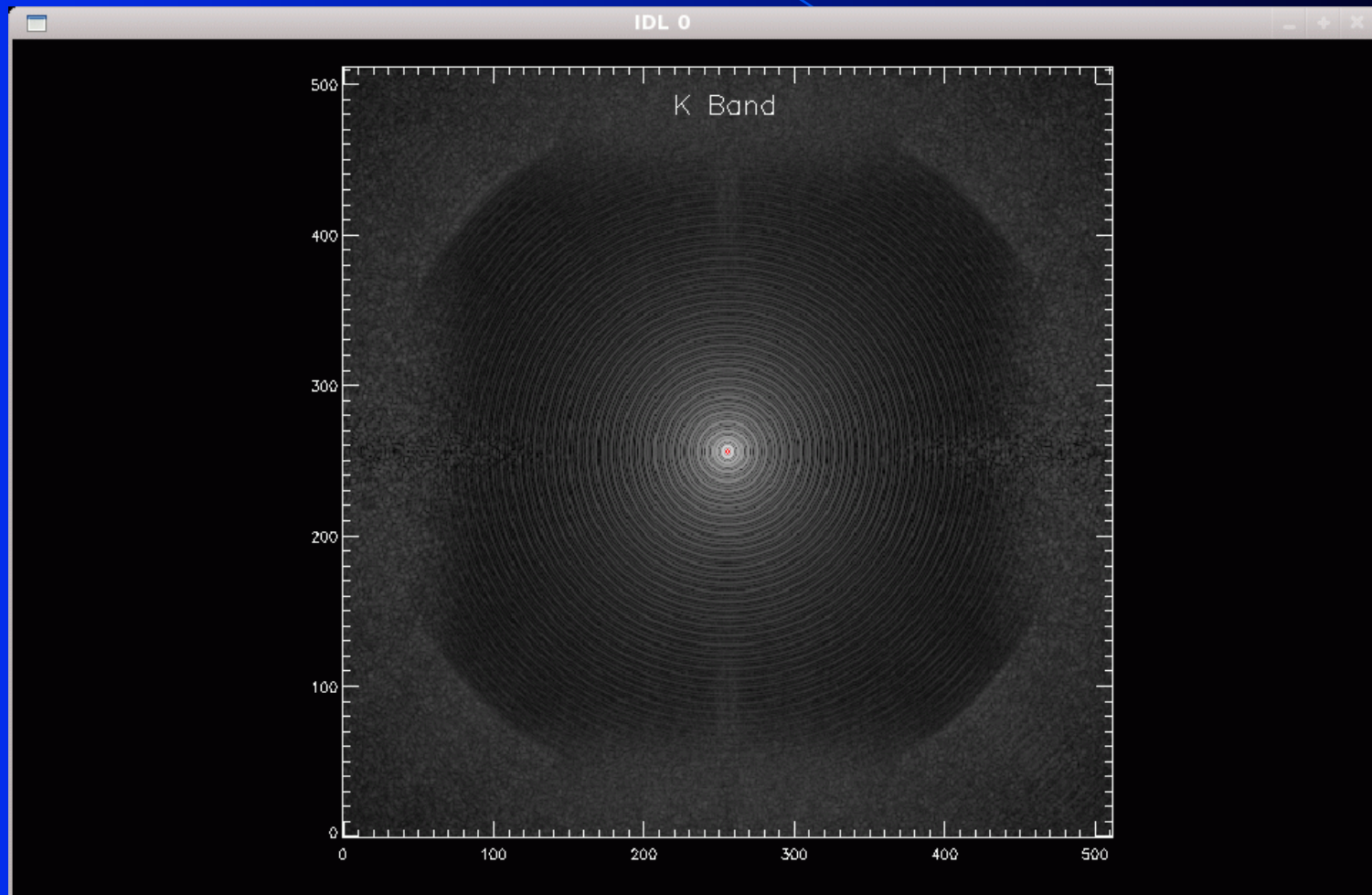


Doesn't seem too problematic,  
although much slower than in typical SH case

# J, H, K PSF comparisons

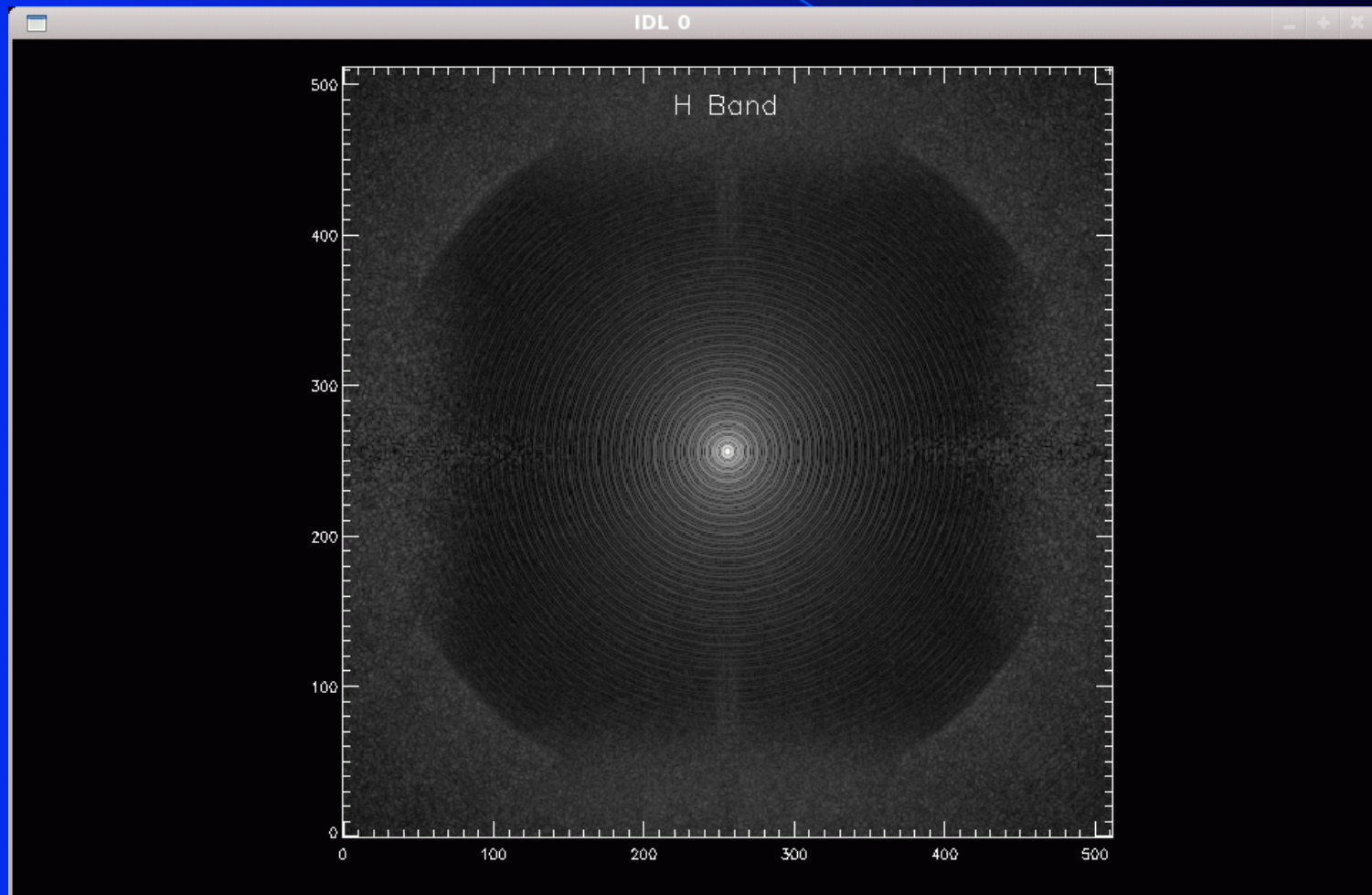


# K Band



Constant sampling in  $\lambda/D$  units

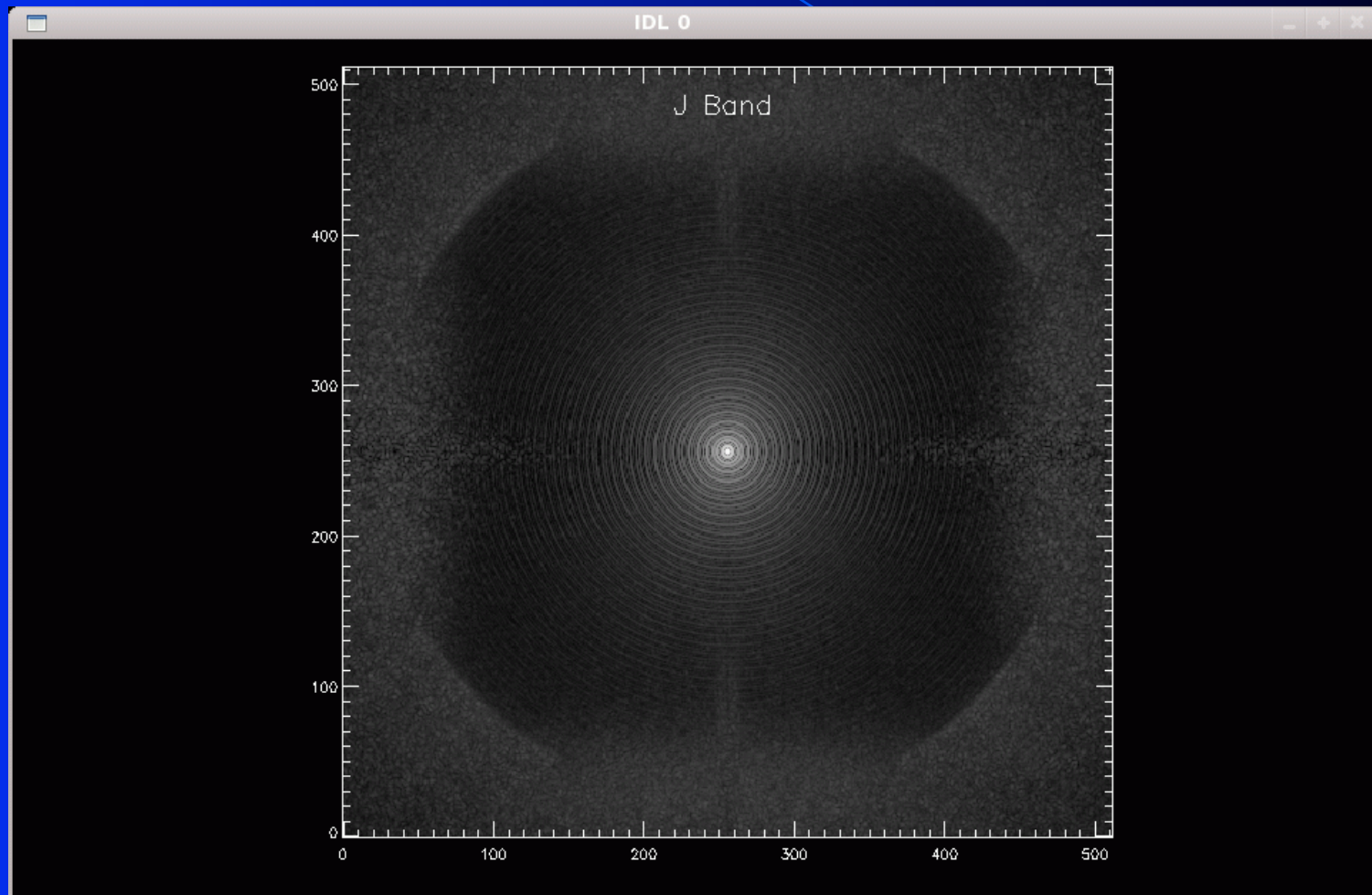
# H Band



Constant sampling in  $\lambda/D$  units



# J Band



Constant sampling in  $\lambda/D$  units

# Comparing Cure(D) and MVM

---

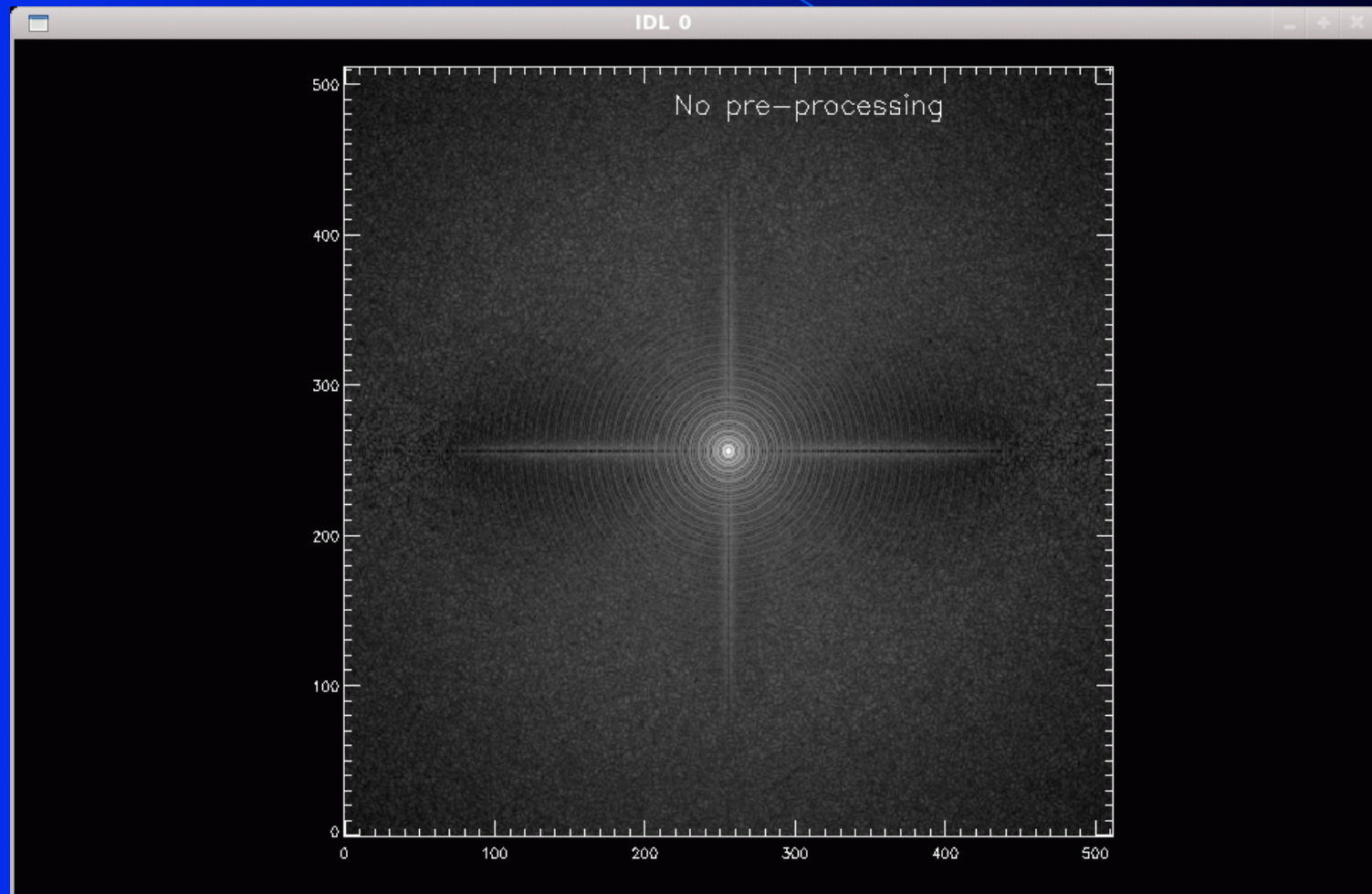
- Idea: See how different 2 reconstruction algorithms are from the PSF point of view.
- Use **same** Measurements to calculate commands:
  - With the standard MVM w/ Interaction matrix inversion (+ some regularization)
  - Cure(D)
- Commands are sent to the **same simulation**, with same input phase screens, noise,...
- Only difference is the reconstructor everything else in the simulation stays the same.

# Cure(D)

---

- Fast reconstruction algorithm developed by the Austrian AO Team (AAO)
- MVM is used as a “reference” case against which Cure(D) is tested.
  - Modal interaction matrix + ad-hoc regularization
- Initial “poor” performance of Cure compared to MVM pushed improvements in Cure → now almost identical performance, BUT many less FLOPs!
- → Shows importance of the performance benchmark
- “*CuRe - Fast wavefront reconstruction algorithm for extremely large telescopes*”, Rosensteiner, M., JOSA A, in press
- “*Cumulative Wavefront Reconstructor for the Shack-Hartman Sensor*”, M. Zhariy, A. Neubauer, M. Rosensteiner, R. Ramlau, Inverse Problems and Imaging, accepted

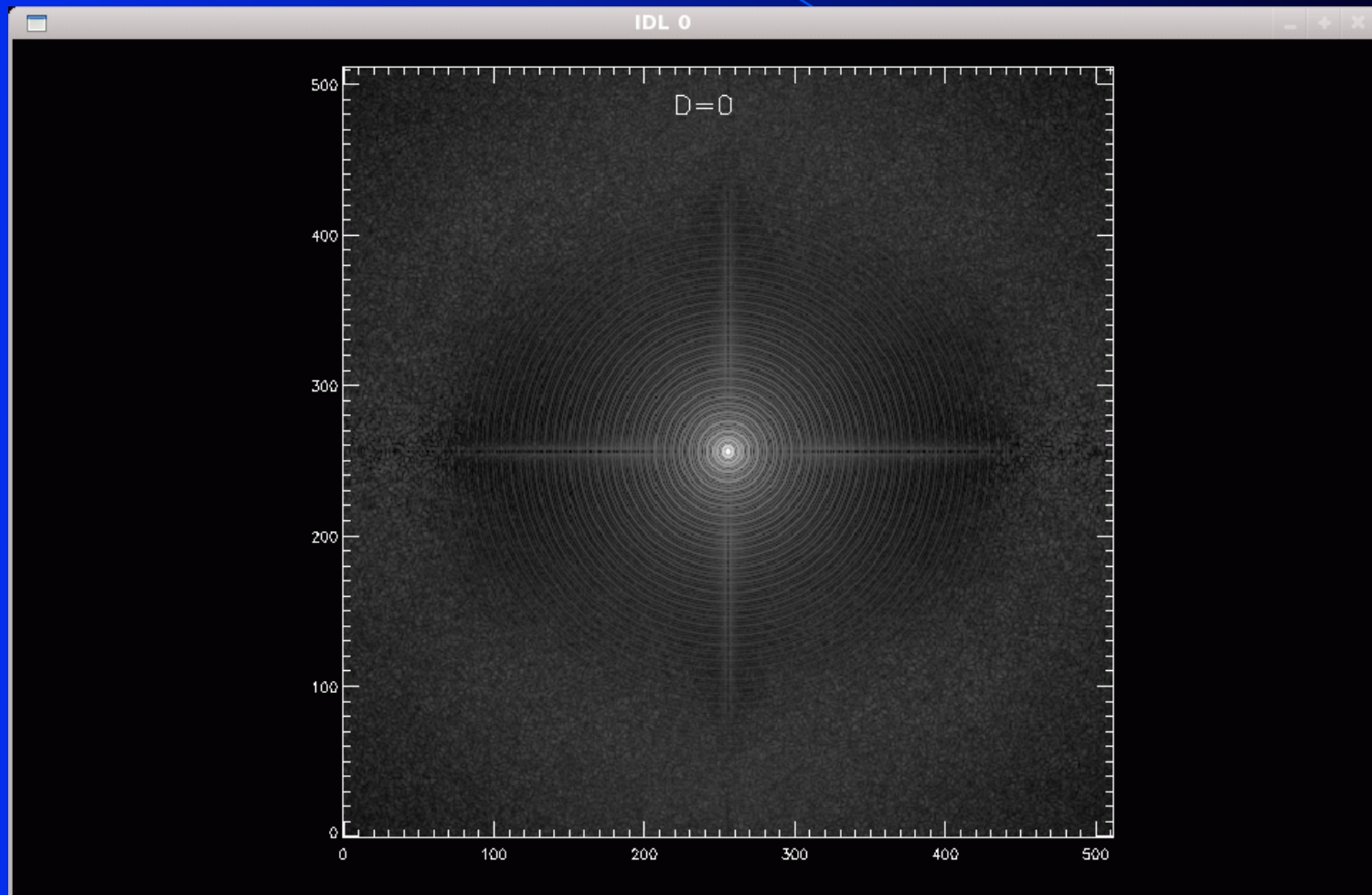
# Cure without pre-processing



with help from the Austrian AO Team

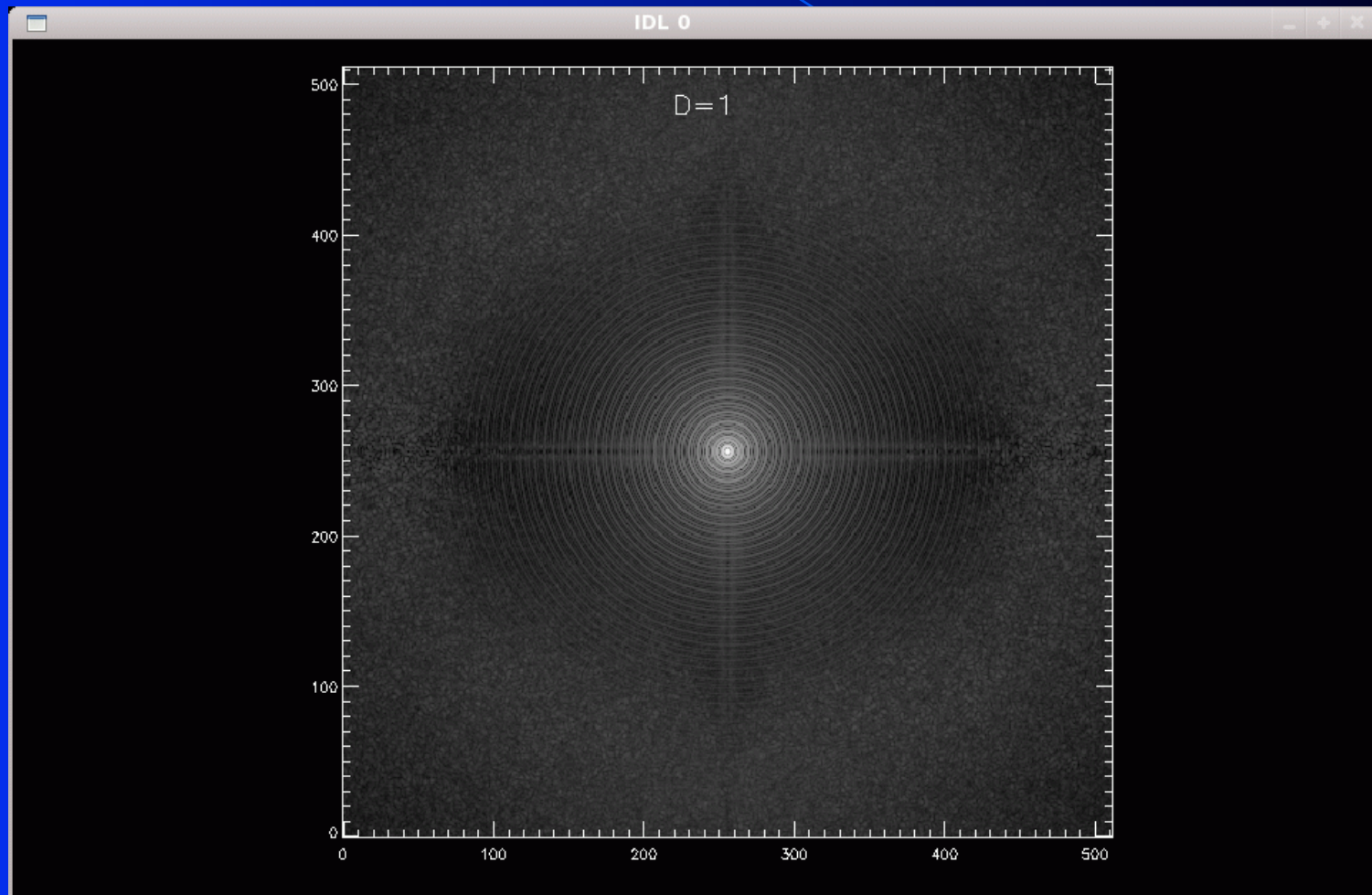


# Cure with pre-processing



with help from the Austrian AO Team

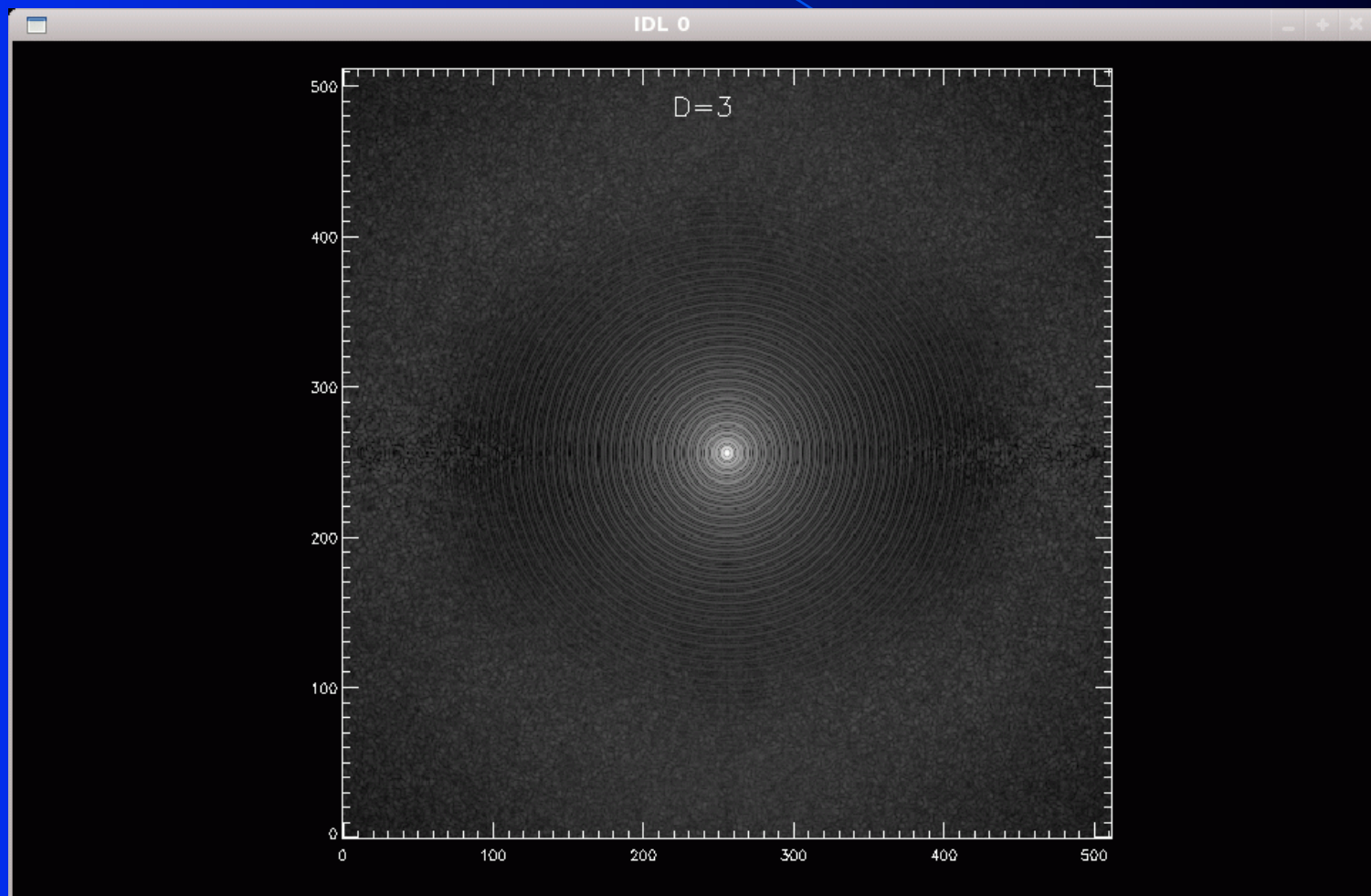
# CureD, D=1 (Domain decomp.)



with help from the Austrian AO Team

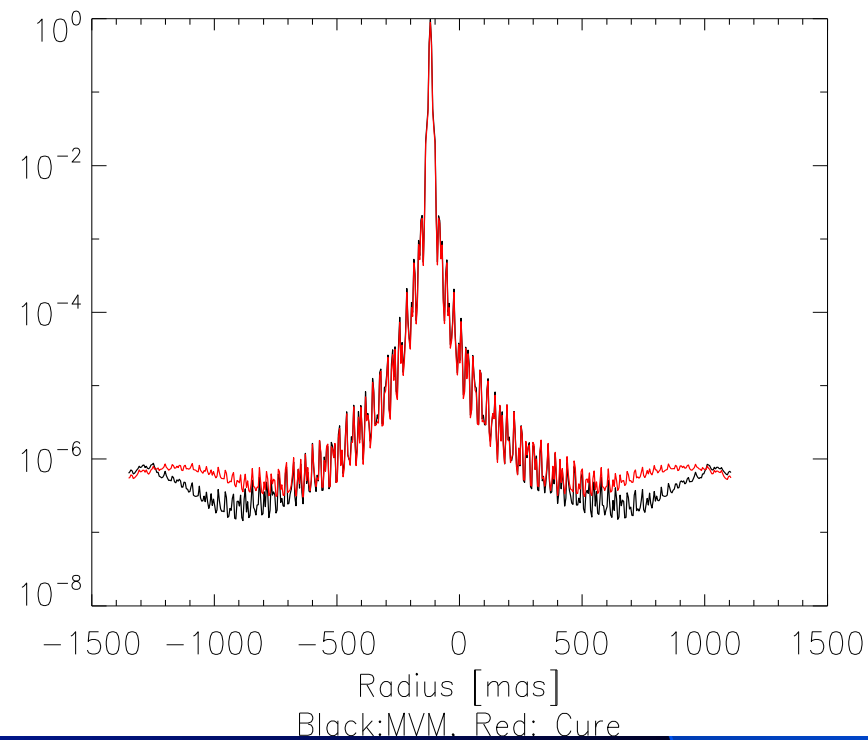
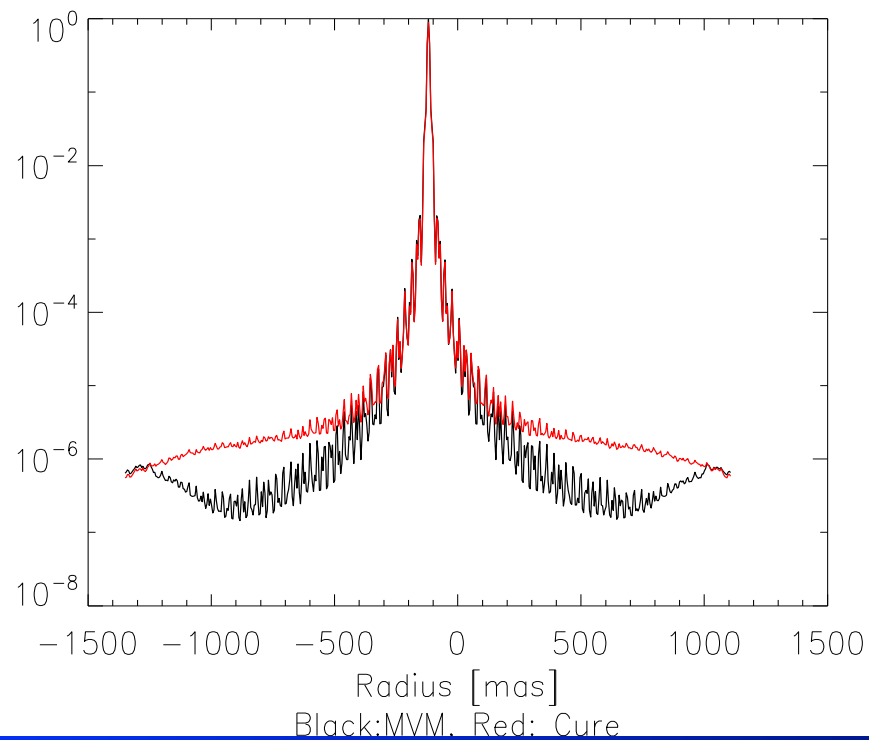


# CureD, $D=3$ (Domain decomp.)



with help from the Austrian AO Team

# Evolution of Cure(D) vs MVM



Very first comparison    After some months of hard work of AAT  
(pre-processing of data, CureD)

# MAORY-like MCAO configuration

---

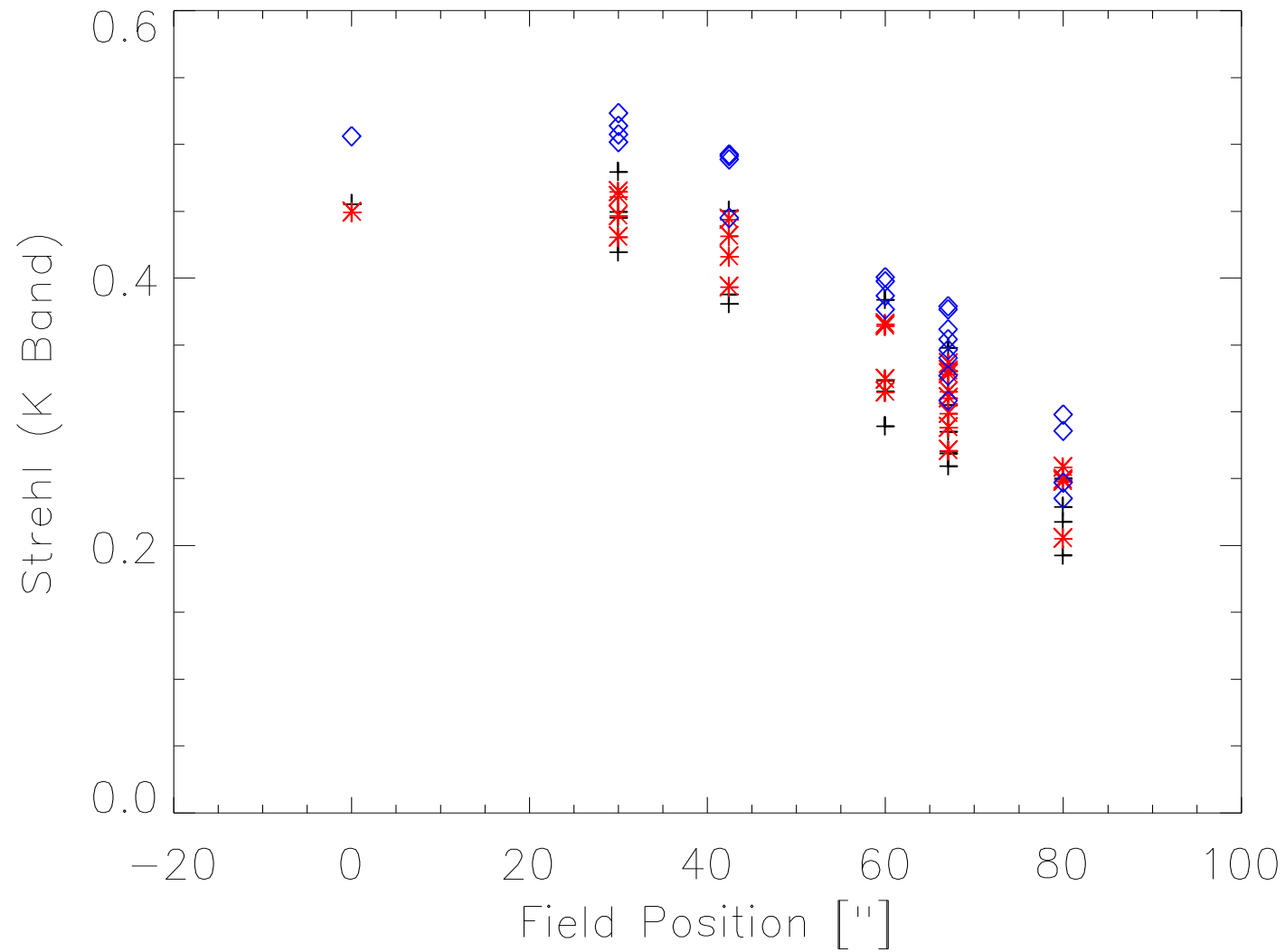
- Continue comparisons of reconstructors
- MCAO is also part of the test cases
- In addition to Cure (MCAO→Kaczmarz), we also test Frim3D
  - “A Kaczmarz type iterative reconstructor for Multi-Conjugate Adaptive Optics”, AO4ELT2, Ramlau & Rosensteiner
  - “Performance of MCAO on the E-ELT using the Fractal Iterative Method for fast atmospheric tomography”, AO4ELT2, Tallon et al.
- 99% same config (still a few discrepancies with NGS sensors, but impact should be minor)

# MAORY-like configuration

---

- 6 Sodium LGS (84x84 WFS)
  - “High flux”
  - Spot elongation neglected for the moment (planned)
  - 90km, fixed
  - 2' (diameter) circle, no central LGS
- 3 NGS
  - one 2x2 (for fast focus + TT)
  - two 1x1 (TT only)
  - “High flux”
- 3 DMs
  - 0 (full) 4km (2 \*spacing) 12.7km (2 \*spacing)
- Corrected FOV: 2.8' (Diameter)
- 25 PSF star measure Strehl (K-band) in FOV
- Seeing: 0.8' '
- 9 layer atmospheric model (none of the reconstructors uses intermediate layers)

# MCAO



MVM:Black, Kaczmarz: Red, Frim: Blue

# MCAO results

---

- Frim3D has best performance (for now ?)
  - Good way to regularize
  - Difference on-axis and close to it
- MVM and Kaczmarz extremely close
- How to improve performance ?
  - Reconstruct more layers and the project on DMs
  - Frim3D: different gain for NGS and LGS
  - [...]
- → Very good results showing consistency, and that gaining significantly on computing power is possible.
- Would be interesting to add other reconstructors to the comparison (MCAO and/or XAO, MOAO, LTAO,...)
  - A beauty contest of reconstructors ?
  - IDL, Matlab, yorick interfaces already exist to use Octopus



# Conclusions

---

- Our simulation software is ready and capable for XAO on the E-ELT
- It is also ready for MCAO on the E-ELT
  - Spot elongation was not considered here, but has been shown to work with MVM and Frim3D.
- Importance of reference points when creating new reconstructors – there are many parameters to tweak, and it's reassuring to have independent methods yielding very similar results
  - Performance is hard to know even with analytic / semi-analytic models (which often lack precisely the parameters you want to optimize)
- Comparisons will continue
  - LTAO, MOAO
  - Spot elongation (already done with Frim, to be done with Cure)
- Other algorithms are being developed by AAO
- Also comparison with RTC implementability (//, pipeline,...).

